

On the Possibility to Synthesize Superheavy Nuclei of Higher Neutron Numbers

Š. Šáro

Department of Nuclear Physics and Biophysics, Faculty of Mathematics, Physics and Informatics,
Comenius University, 842 48 Bratislava, Slovak Republic, e-mail: *saro@fmph.uniba.sk*

Abstract: The laboratory synthesis of nuclei of atomic numbers above $Z = 110$ having the maximum achievable neutron number N is considered. Cold fusion reactions are limited with monotonous decrease of the fusion cross section below 0.01 pb for $Z = 114$. The cross section of hot fusion reactions resulting in the synthesis of nuclei of elements of $Z = 114, 115, 116,$ and 118 show no evidence of the influence of the predicted neutron subshell at $N = 172$. The limitations to create more neutron rich superheavy nuclei is analyzed in the case of radioactive beams with half-lives longer than 2 s. Theoretical and experimental half-life values for nuclei of $Z = 110-118$ are compared and half-life limitations are considered. New attempts to synthesize element of $Z = 120$ and 122 is discussed.

1. Introduction

In the middle of the 1960s nuclei of 12 elements beyond uranium were experimentally synthesized up to lawrencium ($Z = 103$) and rutherfordium ($Z = 104$) due to the stabilizing effect of the nuclear shell structure. At that time the question of the shell structure and its effect on nuclear stability was actually topical and the possibility of the existence of superheavy nuclei (SHN) was also considered [1]. It was proposed that in the case of closed proton and neutron shells the nuclei should have half lives long enough to be experimentally observed. Further theoretical studies led to the most probable closed proton shell at $Z = 114$ and closed neutron shell at $N = 184$ [2–4]. Later some model calculations led to other values of the closed proton shell, at present the dominating values are 114, 120, and 126. The closed neutron shell at $N = 184$ has a relatively stable position in the theory. An important step forward was made in 1967–68 by V.M. Strutinsky [5] presenting quantitative calculation of the microscopic part of the shell correction to the binding energy of heavy nuclei.

The first attempts to synthesize superheavy nuclei around $Z = 114$ were too optimistic. The first calculated half lives were of the order of 10^8-9 years and at that time there were no reliable estimations of fusion reaction cross sections. The available experimental technique gave several basic possibilities for the synthesis: a) fragmentation reactions of two symmetric nuclei with similar proton numbers [7] and “gentle fusion” of two rare earth nuclei [6]; b) complete fusion reactions of the heaviest stable double magic nucleus ^{208}Pb and the neighbouring ^{209}Bi with available stable beams of $Z = 26-34$ [8]. This type of complete fusion reaction led to the synthesis of nuclei of elements of $Z = 107-113$ [9–15]; c) the third possibility was the combination of as heavy transuranium target nuclei as available with beams of light nuclei, giving the desired Z . Yu. Ts. Oganessian realized a se-

ries of experiments of this type of fusion reactions choosing the double magic nucleus of ^{48}Ca as projectile and synthesized the nuclei of elements of $Z = 114, 115, 116,$ and 118 [16–20].

In fact the progress in the synthesis of nuclei with higher and higher Z was influenced by several factors. First of all the availability of the required beams and their intensity delivered from ion sources. The lack of reliable theoretical ideas about the process of fusion of heavy nuclei led to extremely large uncertainties in cross section and half life estimations and from here to many unsuccessful experiments.

2. Cross-section limitation

The production cross section of complete fusion reactions of heavy ions is a result of several physical processes playing role in this process – first of all prompt fission, deep inelastic scattering, complete fusion and compound nucleus survival probability. The process of fusion is influenced also with other effects, like the shell correction energy, nuclear spin, shape of the interacting nuclei and others.

There are several theoretical approaches to calculate the cross section of complete fusion reactions, but in general the reliability of these calculations is problematic. In cold fusion reactions of the double magic target nucleus of ^{208}Pb and the neighbouring ^{209}Bi with neutron rich Ti, Fe, Ni, Zn and similar ions the experimental cross section is steadily decreasing from $5 \cdot 10^{-7}$ barn for $Z = 102$ to $5 \cdot 10^{-14}$ barn for $Z = 113$ [15], i.e. the decrease of Z by one unit results in 6-fold decrease of the cross section in average as is shown in Fig. 1. No measurable influence of the stabilizing shell correction energy, isotopic spin or other parameters on the reaction cross section was observed. No explanation was found for example for the significant jump of the $1n$ channel cross section from $3.3 \left(\begin{smallmatrix} 62 \\ 27 \end{smallmatrix} \right)$ pb for $^{208}\text{Pb}(^{62}\text{Ni}, 1n)$ [12] to $15 \left(\begin{smallmatrix} 9 \\ 6 \end{smallmatrix} \right)$ pb for $^{208}\text{Pb}(^{64}\text{Ni}, 1n)$ [21] differing only by 2 neutrons in the Ni ions.

The hot complete fusion reactions, based on actinoid targets from uranium to californium and on the double magic ^{48}Ca projectile ions has an unexpected feature. In spite of theoretical predictions the cross section of all realized reactions have very similar values differing less than one order of magnitude as it is illustrated in Fig. 2. and Fig. 3. For $Z = 112–118$ and for the neutron number $N = 170–177$ no systematic trend in measured cross section values was observed. It means no observable influence of the predicted neutron subshell at $N = 172$ [23] or the influence of the closed neutron shell at $Z = 184$. To draw serious consequences for the theory, these cross section values need to be confirmed in independent experiments and at higher reliability.

3. Target – projectile limitation

To investigate the region of superheavy elements above $Z = 112$, the only today known approach is the method of hot fusion reaction of actinoid target nuclei and suitable accelerated ions. The successful chain of reactions based on U, Pu, Cm, Bk, and Cf targets and stable ^{48}Ca ions is exhausted. In these reactions the limits of the radioactive target and stable projectile nuclei combinations were reached in the direction to more neutron rich nuclei towards to $N = 184$. To reach higher neutron numbers up to the predicted closed

neutron shell at $N = 184$ actinoid target nuclei have to be bombarded with ions heavier than ^{48}Ca . But even in this case only at $Z = 124$ one can reach $N = 184$ as it is shown in Fig. 4.

In the region of interest ($Z = 114$ – 126) higher neutron numbers are achievable only using radioactive beams (RBs). Considering available actinoid target nuclei and radioactive beams of ions having half lives longer than 2 seconds, the achievable front line is shown in Fig. 4. Even in these case the predicted closed neutron shell at $N = 184$ is achieved at $Z = 119$. Today we have no idea how to reach the proposed closed neutron shall ($N = 184$) at $Z = 114$.

The idea to use radioactive beams to synthesize neutron rich superheavy elements is not a new one. The basic problem is the available intensity of such beams. Suitable neutron rich radioactive ions can be created by fragmentation of high energy nuclei and by consequent in-flight separation and deceleration of the separated high energy ion to coulomb barrier energy level. With respect to the expected picobarn (10^{-40} m^2) cross section level, the necessary beam intensity is of the order of $1 \text{ p}\mu\text{A}$ or 10^{12-13} ions/s. The present approachable intensity of single ion RBs is below 10^9 ions/s. To reach the 10^{12-13} ions/s level will be a very difficult task. New powerful ion source of primary beams, high energy accelerator to accept such beams, and probably the parameters of the fragment separator and decelerating ring should be specially adjusted to fulfill the task.

4. Half life limitation

The first phase of a heavy compound nucleus creation in a complete fusion reaction is a complex function of many internal and external parameters of both interacting nuclei. After the formation of the compound nucleus its survival probability depends on fewer internal parameters of the formatted compound nucleus itself. This gives for the theory the possibility to calculate more reliable half-life values than in the case of the fusion probability.

Half life calculations made by Sobiczewski et al. [24] show a clear dependence of alpha decay half lives on both, the proton number Z and the neutron number N of a particular nucleus. With increasing atomic number Z the half lives of all isotopes of the given Z are monotonously decreasing, but there are two significant peaks at neutron numbers around the proposed neutron subshell of $N = 162$ and closed neutron shell of $N = 184$ (see Fig. 5).

To check the reliability of these calculations for superheavy elements is rather problematic. First of all the theory gives data only for even-even nuclei, but in the matrix of 121 nuclei of $Z = 110$ – 120 and $N = 166$ – 176 only 9 even-even nuclei have experimentally determined alpha-decay half lives. The comparison of the available experimental and calculated alpha decay half lives are given in Tab. 1. The calculated values are mostly underestimated and differ from the experimental ones from several times to two orders of magnitude.

The second problem is the uncertainty in the experimental data due to very low statistics, in some cases only one or two recorded events.

5. Present attempts

At present new attempts are made to go further in SHN synthesis. The synthesis of $Z = 120$ element is on the program in two laboratories. In JINR Dubna the reaction of $^{244}\text{Pu} + ^{58}\text{Fe} \rightarrow ^{299}\text{120} + 3\text{n}$ is going on [22] and at GSI Darmstadt [25] the reaction of $^{238}\text{U} + ^{64}\text{Ni} \rightarrow ^{299}\text{120} + 3\text{n}$ is on the way. Both reactions are leading to the same compound nucleus $^{302}\text{120}^*$ and the 3n evaporation channel will create the same evaporation residue of $^{299}\text{120}$. The expected alpha decay chain of $^{299}\text{120}$ after one unknown member ($^{295}\text{118}$) will follow the path of already synthesized alpha decay nuclei - $^{291}\text{116}$ - $^{287}\text{114}$ - $^{283}\text{112}$ - $^{279}\text{110}$ as it is illustrated in Fig. 6.

The cross section of both reactions leading to $^{299}\text{120}$ is uncertain. The hot fusion synthesis of all nuclei of elements of $Z = 112, 114, 116,$ and 118 were based on the interaction of the neutron rich double magic ^{48}Ca ions with transuranium target nuclei. If the closed shells in ^{48}Ca played a substantial stabilizing role in the process of fusion then the cross section of both reactions, leading to $^{299}\text{120}$ can fall significantly below 1 pb. The calculated alpha decay half life of $^{299}\text{120}$ is about 1 μs [24] or higher [26]. The time of flight of the evaporation residues from the target to the analyzing detector array is several μs , therefore the detection efficiency may be critical.

6. Perspectives

The cross section of the cold fusion reaction at $Z = 113$ is only 0.05 pb [15]. The monotonous decrease of σ from $Z = 102$ to $Z = 113$ predicts the expected value of σ for $Z = 114$ 0.01 pb. This is below the acceptable beam time of several months to synthesize one nucleus of element $Z = 114$. There are expectations to design ECR ion sources delivering heavy ion beams of the order of 10^{14} ions/s which should allow to reach the 0.01 pb level at reasonable beam time. But at such a heavy ion beam intensity several secondary problems will appear. First of all the energy deposition in Pb or Bi targets is limiting at present the acceptable beam intensity to about 10^{12} ions/s. The second problem will be the background. The method of identification of new nuclei based on the alpha - alpha correlation method requires as low background count rate in the sensitive part of the energy spectra as possible. To avoid difficulties of this type special effort should be paid to the construction and shielding of all parts of the equipment from the target chamber to the detector array beyond the separator.

Properly designed transuranium targets will be able to accept beams of ions of the order of 10^{14} but the problem of the background will be serious also here, especially in the case of gas filled separators.

7. Conclusion

The artificial synthesis of nuclei heavier than uranium was, all the time, a front-line experiment requiring the most advanced laboratory equipment and novel physical approaches. The method used to synthesize the first transuranium nuclei exhausted its possibilities at mendelevium Md ($Z = 101$). The next generation of experiments postponed the frontier to siborgium ($Z = 106$). A new approach was needed to continue, the

cold fusion concept, which shifted the frontier to element 113, where the cross section of the 1n evaporation channel fall to 0.05 pb. An unexpected success of hot fusion reactions, based of the double magic ^{48}Ca ions and trasuranium targets pushed the frontier to element 118. The present attempts to go further are not based on novel physical ideas but on the expectation that the 3n hot fusion evaporation channel will work as well as in the case of the double magic ^{48}Ca ions. To get closer to the predicted closed proton shell at $Z = 114$, 120, or 126 and closed neutron shell at $N = 184$ new physical ideas and more advanced and powerful experimental technique should be involved.

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Table 1. Measured (upper data) and calculated (Sobiczewski et al. [24]) half-lives of the synthesized elements of $Z = 110 - 118$.

120						?
						1 - 2 ms
118						0.9 ms 0.05 ms
116					15 ms	18 ms 8 ms
114				160 ms 8 ms	800 ms 200 ms	
112			0.50 ms 7 ms	101 ms 1 ms		
110	9.7 ms 0.2 ms					
Z / N	166	168	170	172	174	176

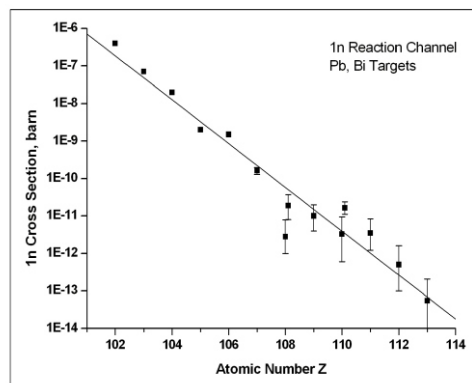


Fig. 1. Experimental cross section for cold fusion reactions for evaporation residues of $Z = 102 - 113$ and 1n evaporation channel.

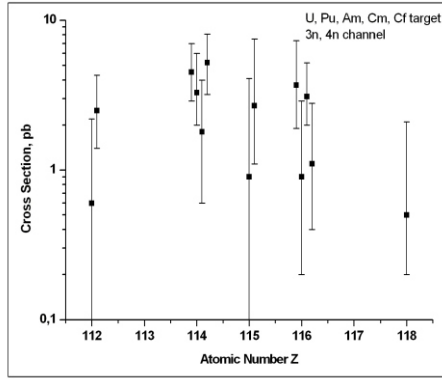


Fig. 2. Experimental cross section for hot fusion reactions for evaporation residues of $Z = 112-118$ and $3n, 4n$ evaporation channels [16-20].

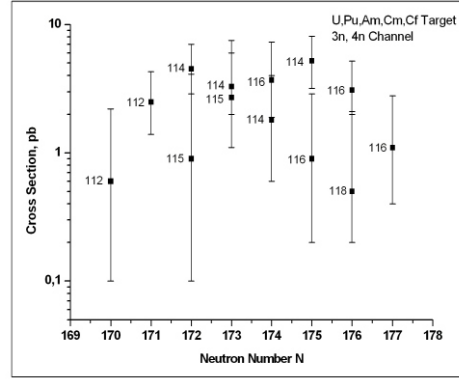


Fig. 3. Experimental cross section for hot fusion reactions for evaporation residues of $N = 170-177$ and $Z = 112-118$ [16-20].

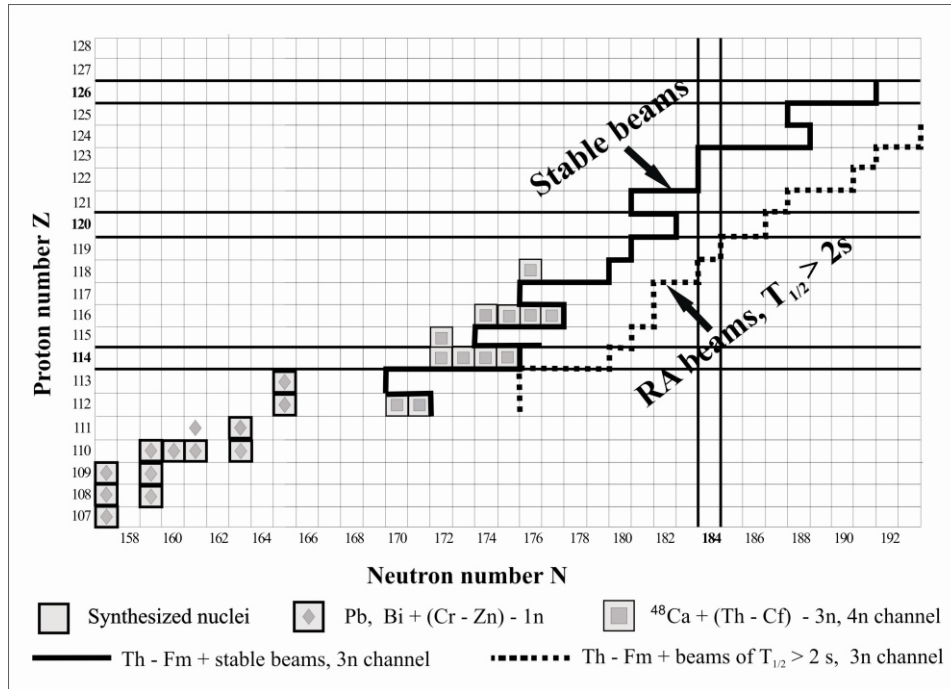


Fig. 4. Chart of superheavy nuclei. The full line represents the maximum possible number of neutrons in the particular nuclei, created in $3n$ evaporation channel hot fusion reactions of actinoid target nuclei and stable projectiles. The dashed line represents the same, but for radioactive projectiles having half-lives $>2\text{s}$.

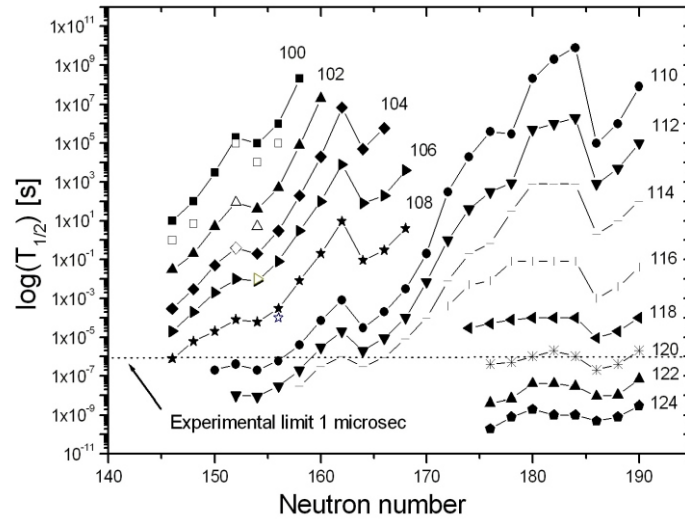


Fig. 5. Alpha decay half life calculation for isotopes of heavy elements from fermium (Z = 100) to element of Z = 124 (after Sobiczewski et al [24]).

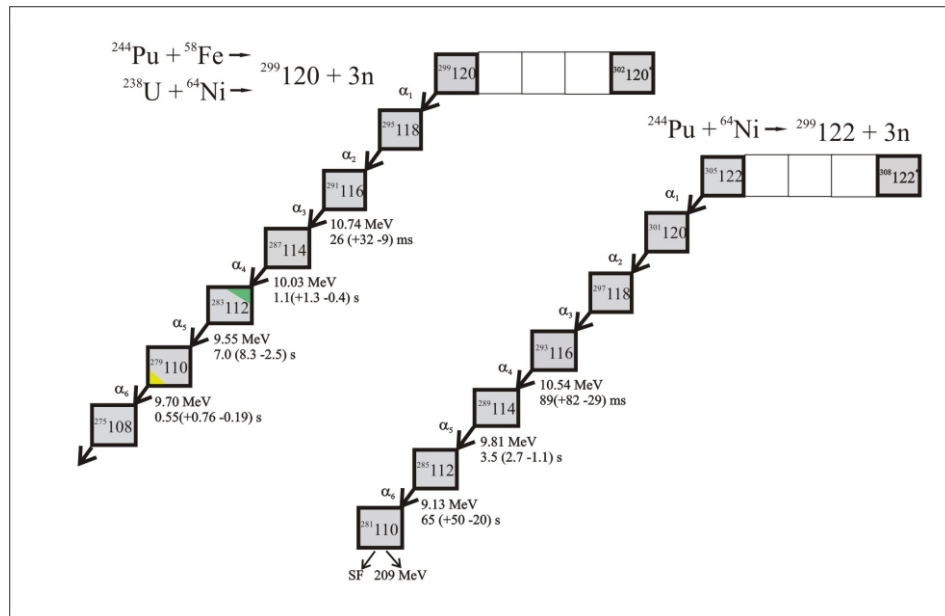


Fig. 6. The expected decay chains of evaporation residues $^{299}_{120}$ and $^{305}_{122}$.

References

- [1] W. D. Myers, W. J. Swiatecki: Nuclear Phys. **81** (1966) 1.
- [2] A. Sobiczewski, F. A. Gareev, B. N. Kalinkin: Phys. Lett. **22** (1966) 500.
- [3] H. Meldner: Ark. Fys. **36** (1967) 593.
- [4] U. Mosel, W. Greiner: Z. Phys. **222** (1969) 261.
- [5] V. M. Strutinsky: Nuclear Phys. A **95** (1967) 420; Nuclear Phys. A **122** (1968) 1.
- [6] W. Nörenberg: GSI-95-08 Preprint. Januar 1995.
- [7] A. Sandulescu et al.: Physical Letters, **60B** 3 (1976) 225.
- [8] A. S. Ilionov, Yu. Ts. Oganessyan, E. A. Cherepanov: Sov. J. Nucl. Phys. **36**(1) (1982) 69.
- [9] G. Münzenberg et al.: Z. Phys. **A300** (1981) 107.
- [10] G. Münzenberg et al.: Z. Phys. **A309** (1982) 89; Z. Phys **A315** (1984) 145.
- [11] G. Münzenberg et al.: Z. Phys. **A317** (1984) 235; Z. Phys **A328** (1987) 49.
- [12] S. Hofmann et al.: Z. Phys. **A350** (1995) 277 280; Eur. Phys. J. **A10** (2001) 5 10.
- [13] S. Hofmann et al.: Z. Phys. **A350** (1995) 281 282; Eur. Phys. J. **A14** (2002) 147 157.
- [14] S. Hofmann et al.: Z. Phys. **A354** (1996) 229 280; Eur. Phys. J. **A14** (2002) 147 157.
- [15] K. Morita et al.: J. Phys. Soc. Jpn. **73** (2004) 2593.
- [16] Yu. Ts. Oganessian et al.: Eur. Phys. J. A **15** (2002) 201 204.
- [17] Yu. Ts. Oganessian et al.: Phys. Rev. C **69**, 021601(R) (2004).
- [18] Yu. Ts. Oganessian et al.: Phys. Rev. C **69**, 054607 (2004).
- [19] Yu. Ts. Oganessian et al.: Phys. Rev. C **70**, 064609 (2004) RAPID IONS.
- [20] Yu. Ts. Oganessian et al.: Phys. Rev. C **74**, 044602 (2006) RAPID IONS.
- [21] S. Hofmann et al.: Proc. CP425, Tours Symposium on Nuclear Physics III, edited by M. Arnould et al., 1998, The American Institute of Physics 1-56396749-9/98.
- [22] M. G. Itkis: Report on the 101st session of the JINR Scientific Council, January 19 20, 2007.
- [23] W. Zhang, J. Meng, S. Q. Zhang, L. S. Geng, H. Toki: Nuclear Phys. A **753** (2005) 106.
- [24] A. Sobiczewski, K. Pomorski: Progress in Particle and Nuclear Physics **58** (2007) 292 349.
- [25] S. Hofmann et al.: to be published.
- [26] J.-F. Berger, L. Bitaud, J. Dechargé, M. Girod, K. Dietrich: Nuclear Phys. A **685** (2001)1c.